Light and Charge Measurements of Simulated Aluminum Micrometeoroids

Michael DeLuca^{1,2,3}, Evan Thomas^{2,4}, Tobin Munsat^{2,4}, Robert Marshall³, Zoltan Sternovsky^{1,2,3}

¹ LASP, CU Boulder ² IMPACT, CU Boulder

³ Aerospace Engineering, CU Boulder

⁴ Physics, CU Boulder



Laboratory for Atmospheric and Space Physics University of Colorado **Boulder**

Meteors in Earth and Space Science

- Atmospheric and Oceanic Effects
 - Deposit metals such as Fe, Na, K in upper atmosphere
 - Produces metal layers in atmosphere
 - Consequences for models of upper-atmospheric dynamics
 - Bio-available Fe filters down into the ocean



Image from earthobservatory.nasa.gov



- Dust in the Solar System
 - Models of source bodies (comets and asteroids)
 - Meteoroid distribution in the near-Earth environment
 - Poses danger to spacecraft

Image from apod.nasa.gov

Problem: What is the total mass of cosmic dust entering Earth's atmosphere from space?

- Mass input estimates range from 5 to 270 metric tons/day (Plane, 2012)
- Most of this mass is in small particles that ablate in the mesosphere
 - Mass distribution peaks at 10⁻⁵ g
 - Radars are sensitive to 10^{-9} g 10^{-3} g



Ablation Experiments



Image from impact.colorado.edu



Particles shot into ablation chamber: Speed: 10-70 km/s Type: Fe or Al Radius: fraction of a micron $Gas: N_2$, air, O_2 , or CO_2 at p = 0.01 – 0.5 Torr

Figure by J. Simolka (2013)

Ionization Coefficient β

 β =number of electrons/number of ablated atoms = $(\Delta Q/e)/(\Delta m/m\downarrow atom)$



 $\beta \downarrow 0 = c(\nu - \nu \downarrow 0) \uparrow 2 \nu \uparrow 0.8 / 1 + c(\nu - \nu \downarrow 0) \uparrow 2 \nu \uparrow 0.8$

 $\beta(v) = \beta \downarrow 0 \ (v) + (1 - \beta \downarrow 0 \ (v))(1 + \mu)^2 \ /2v^2 \mu \int v \downarrow 0 \ (v) \# \beta(v^*) v^* dv$

Past: Iron Ablation Experiments

• Charge measurements for Fe dust particles in CO₂, O₂, N₂, air



Present: Aluminum Ablation Experiments

• Charge measurements for Al dust in air



Aluminum Beta Results



Fit to Power Law gives: $\beta = 4.491 \times 10^{7} - 5 \nu^{2} \cdot 131$

Fit to Jones gives: $c=6.6911 \times 10^{7}-6$ ($c=10.8320 \times 10^{7}-6$ if neglect integral term)

Threshold Velocity for Al: $v \downarrow 0 = 9.107 \text{ km/s}$

Light Measurements

• Four 16-Channel PMT's





Hamamatsu R5900U-16-L20



Particle Tracking







PMT Channels for Exp 10, Run 17, Event 15 (p = 90 mtorr, $r_0 = 103$ nm)



PMT Channels for Exp 10, Run 17, Event 183 (p = 90 mtorr, $r_0 = 125$ nm)





PMT Velocity Fit for Exp 10, Run 17, Event 15 (p = 90 mtorr, $r_0 = 103$ nm)



PMT Velocity Fit for Exp 10, Run 17, Event 183 (p = 90 mtorr, r₀ = 125 nm) 0.35 • • • • • PMT Pulses • • • PMT V = 14.7003 kps





PMT Channels for Exp 10, Run 22, Event 24 (p = 70 mtorr, r₀ = 47 nm)



PMT Channels for Exp 10, Run 24, Event 65 (p = 100 mtorr, $r_0 = 104$ nm)



PMT Velocity Fit for Exp 10, Run 19, Event 78 (p = 200 mtorr, $r_0 = 177$ nm) 0.35 PMT Pulses PMT V = 10.7367 kps 0.30 Entry V = 12.1969 kps Ê 0.25 ⊧ ge ' _{0.20} ' 0.15 0.10 25 10 15 20 Time (us)





 $\label{eq:time} Time~(us)$ PMT Velocity Fit for Exp 10, Run 24, Event 65 (p = 100 mtorr, r_{0} = 104 nm)



Summary and Next Steps

- Ablation Experiments with Aluminum
 - Determined β for particles from 10.8 km/s to 73.4 km/s
 - Fit β to Power Law and Jones Curve
 - Tracked Particle using PMT's
- Next Steps:
 - Measure slowdown with a new pickup-tube detector
 - Model the ablation of Aluminum micrometeoroids



References

- Jones, W., Theoretical and Observational Determinations of the Ionization Coefficient of Meteors. *Mon. Not. R. Astron. Soc.* 288, 995-1003 (1997).
- Plane, J.M.C., Cosmic Dust in the Earth's Atmosphere. *Chem. Soc. Rev.* 41, 6507-6518 (2012).
- Simolka, J., Design, Manufacture, and Testing of an Experimental Set-Up Simulating Meteoric Ablation in Laboratory Conditions. Diploma Thesis (2013).
- Thomas, E., et al, Measurements of the Ionization Coefficient of Simulated Iron Micrometeoroids. *Geophys. Res. Lett.* 43, 3645-3652 (2016).
- Vondrak, T., et al, A Chemical Model of Meteoric Ablation. Atmos. Chem. Phys. 8, 7015-7031 (2008).